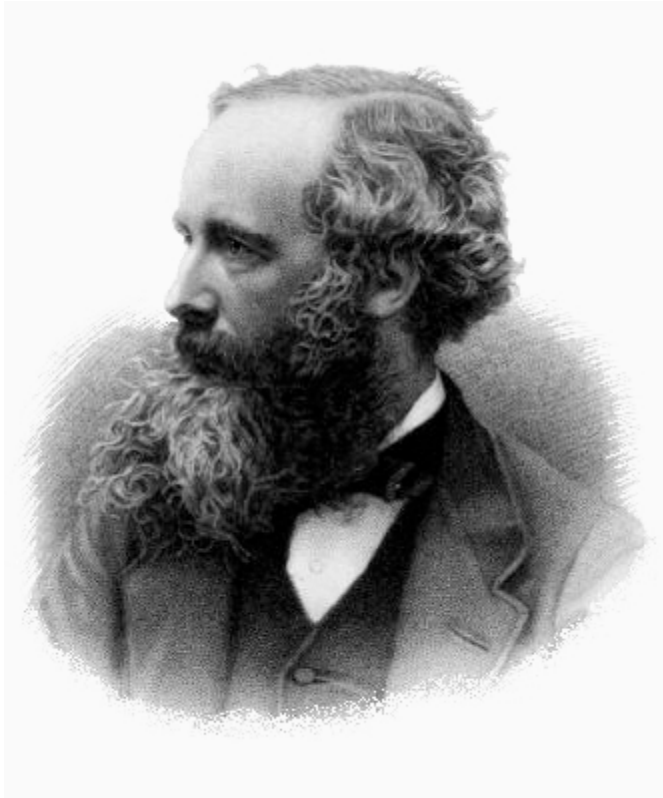
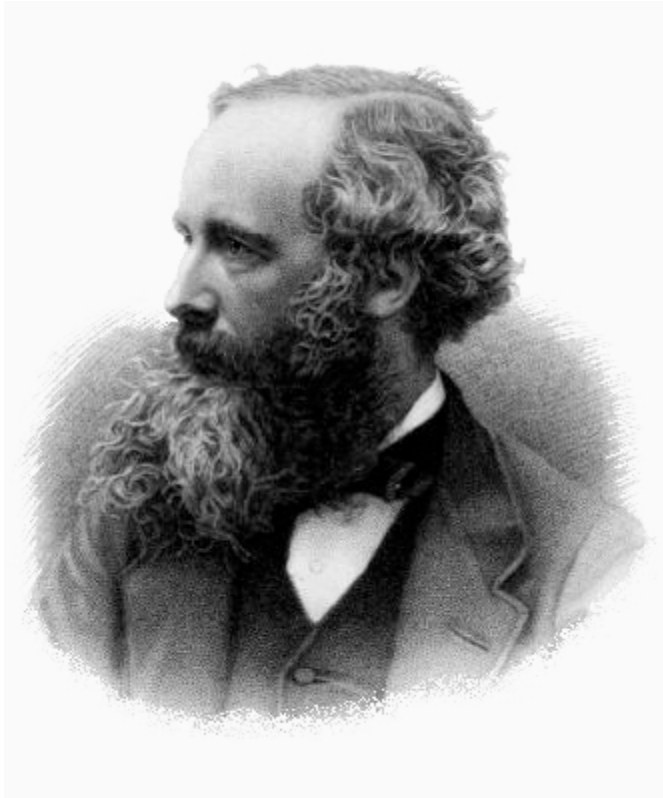


# The Death of Classical Physics



## The Rise of the Photon

# A fundamental question: What is Light?



James Clerk Maxwell  
1831-1879

Electromagnetic Wave

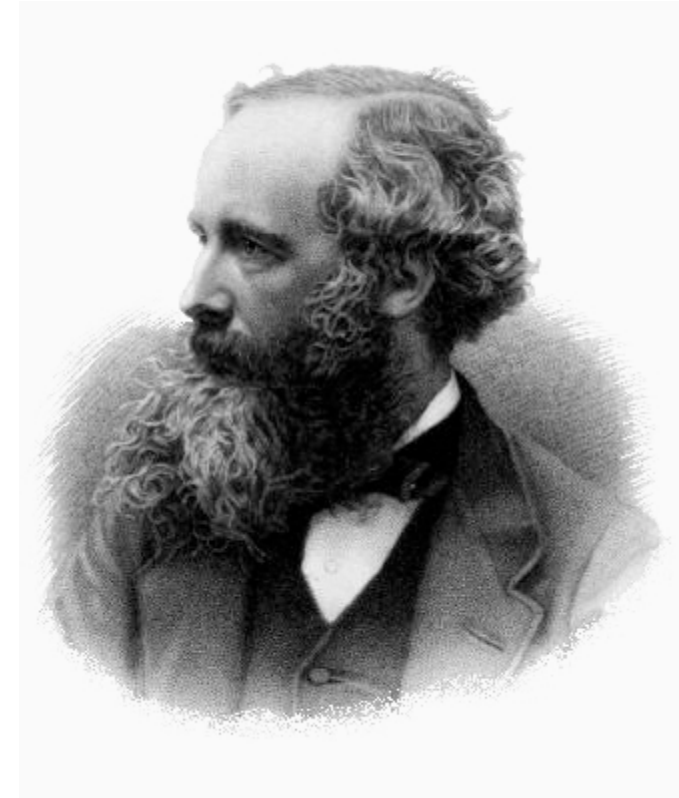


Max Planck  
1858-1947

Photon

# Maxwell's Equations (1865)

Maxwell's equations are a set of four partial differential equations that describe the properties of the electric and magnetic fields and relate them to their sources, charge density and current density.



James Clerk Maxwell  
1831-1879

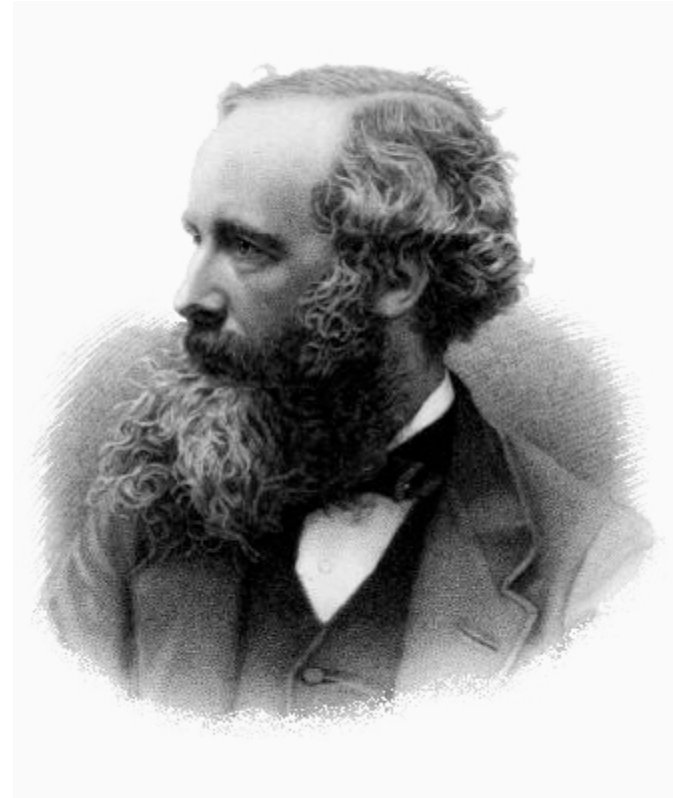
# Maxwell's Equations (1865)

$$\nabla \cdot \mathbf{E} = \frac{\rho}{\epsilon_0}$$

$$\nabla \cdot \mathbf{B} = 0$$

$$\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t}$$

$$\nabla \times \mathbf{B} = \mu_0 \mathbf{J} + \mu_0 \epsilon_0 \frac{\partial \mathbf{E}}{\partial t}$$



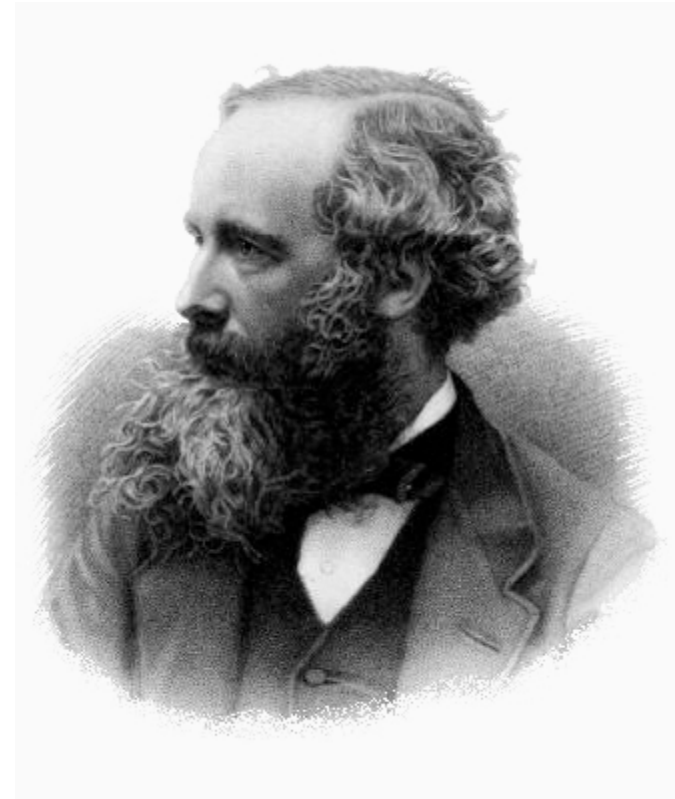
James Clerk Maxwell  
1831-1879

# Electromagnetic Waves - Theory (1865)

$$\nabla^2 \mathbf{E} = \mu_0 \epsilon_0 \frac{\partial^2 \mathbf{E}}{\partial t^2} = \frac{1}{c^2} \frac{\partial^2 \mathbf{E}}{\partial t^2}$$

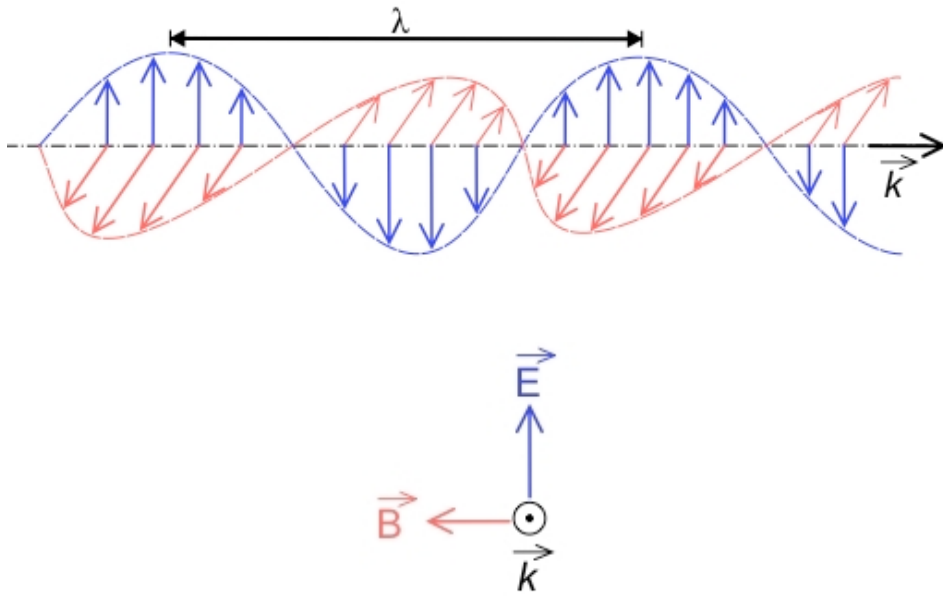
$$\nabla^2 \mathbf{B} = \mu_0 \epsilon_0 \frac{\partial^2 \mathbf{B}}{\partial t^2} = \frac{1}{c^2} \frac{\partial^2 \mathbf{B}}{\partial t^2}$$

$$c = \frac{1}{\sqrt{\mu_0 \epsilon_0}}$$



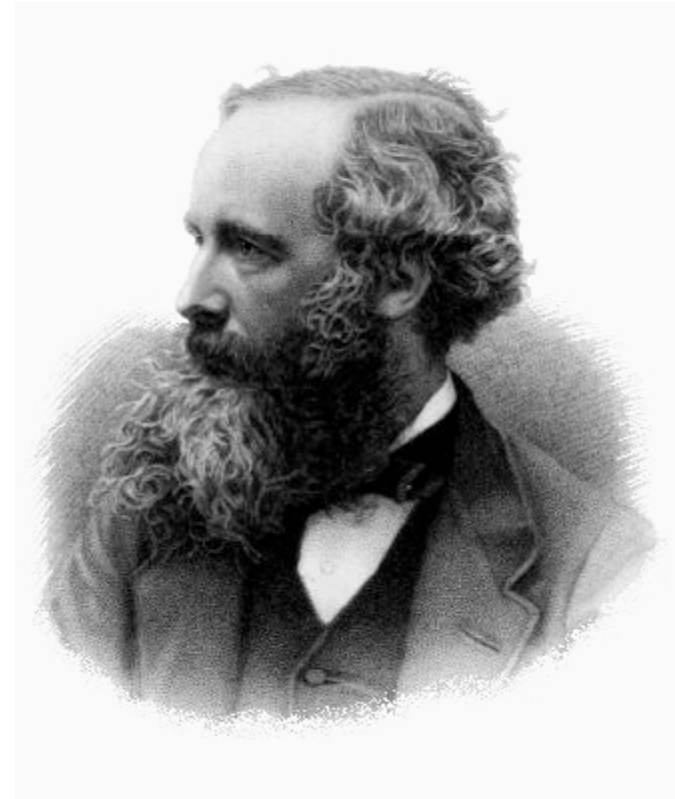
James Clerk Maxwell  
1831-1879

# Electromagnetic Waves - Theory (1865)



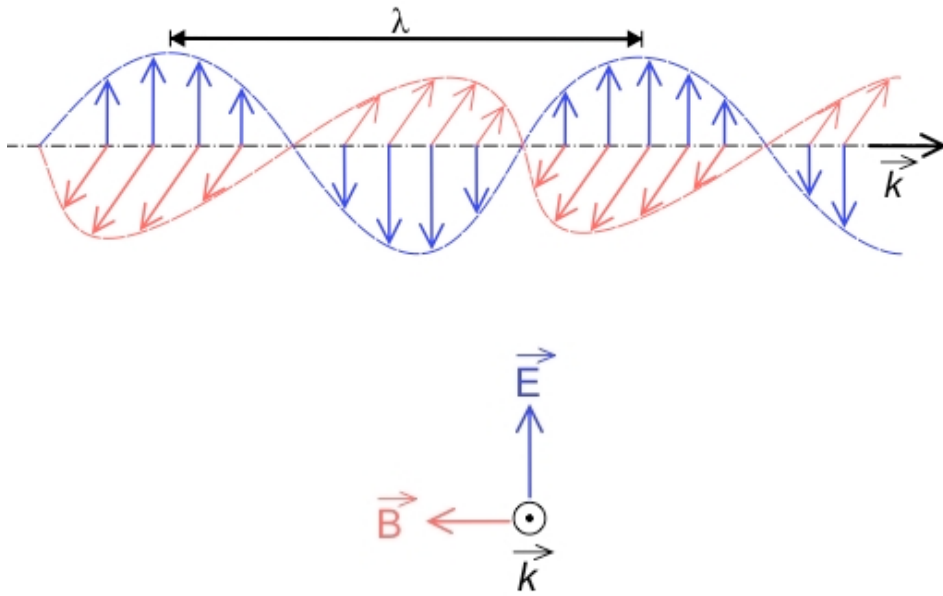
**E and B are "fields" that  
vary with time and space**

(k is the direction of the light wave)



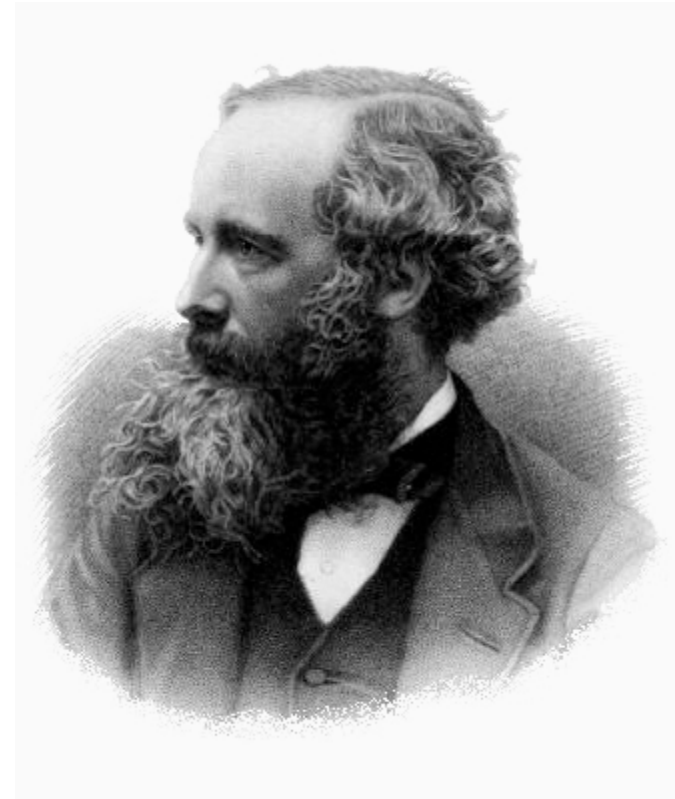
James Clerk Maxwell  
1831-1879

# Electromagnetic Waves - Theory (1865)



$$E(x,t) = E_0 \sin\left[\frac{2\pi}{\lambda}(x - ct)\right]$$

traveling wave equation in one dimension



James Clerk Maxwell  
1831-1879

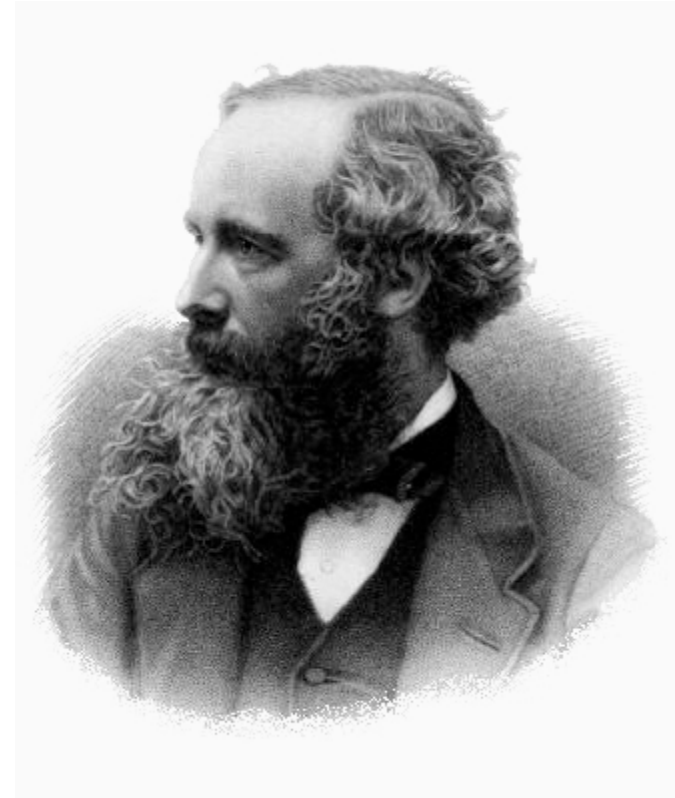


# Electromagnetic Waves - Theory (1865)

$$c = \frac{1}{\sqrt{\mu_0 \epsilon_0}}$$

299,792,458 m/s

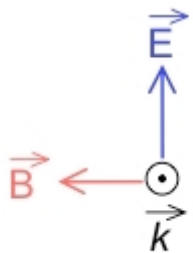
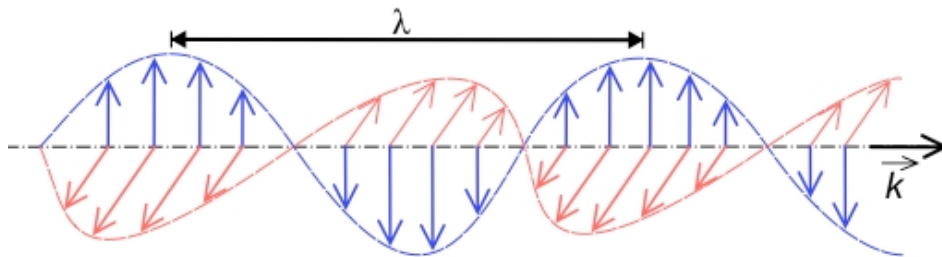
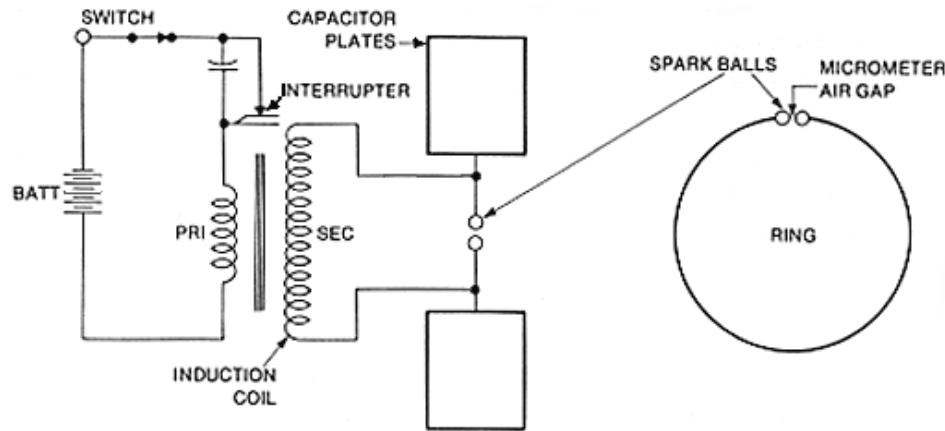
Electromagnetic Waves  
are Light!!!



James Clerk Maxwell  
1831-1879



# Electromagnetic Waves - Experiment (1887)



E and B are fields that  
act on charged particles



Heinrich Hertz  
1857-1894

# Electromagnetic Waves - Experiment (1887)

$$c = \frac{1}{\sqrt{\mu_0 \epsilon_0}}$$

299,792,458 m/s

Electromagnetic Waves  
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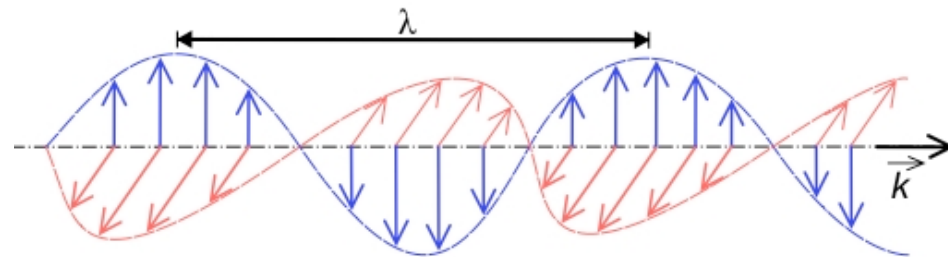
Heinrich Hertz  
1857-1894

# The Electromagnetic Spectrum

$$c = \lambda \nu$$

$$2.998 \times 10^8 \text{ meters/sec}$$

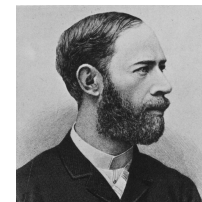
$\lambda$  = the Greek letter lambda  
 $\nu$  = the Greek letter nu



velocity = wavelength x frequency

$$\text{meters/sec} = (\text{meters}) \times (\text{sec})^{-1}$$

The units  $(\text{sec})^{-1}$  are called Hertz (Hz)!



## An example calculation of frequency

**Question:** Calculate the frequency of “typical” red light with a wavelength of 600.0 nm.

$$c = 2.998 \times 10^8 \text{ m/s} = \lambda \nu$$

$$\nu = (2.998 \times 10^8 \text{ m/s}) / (600.0 \times 10^{-9} \text{ m})$$
$$4.997 \times 10^{14} \text{ s}^{-1} \text{ or } 4.997 \times 10^{14} \text{ Hz.}$$

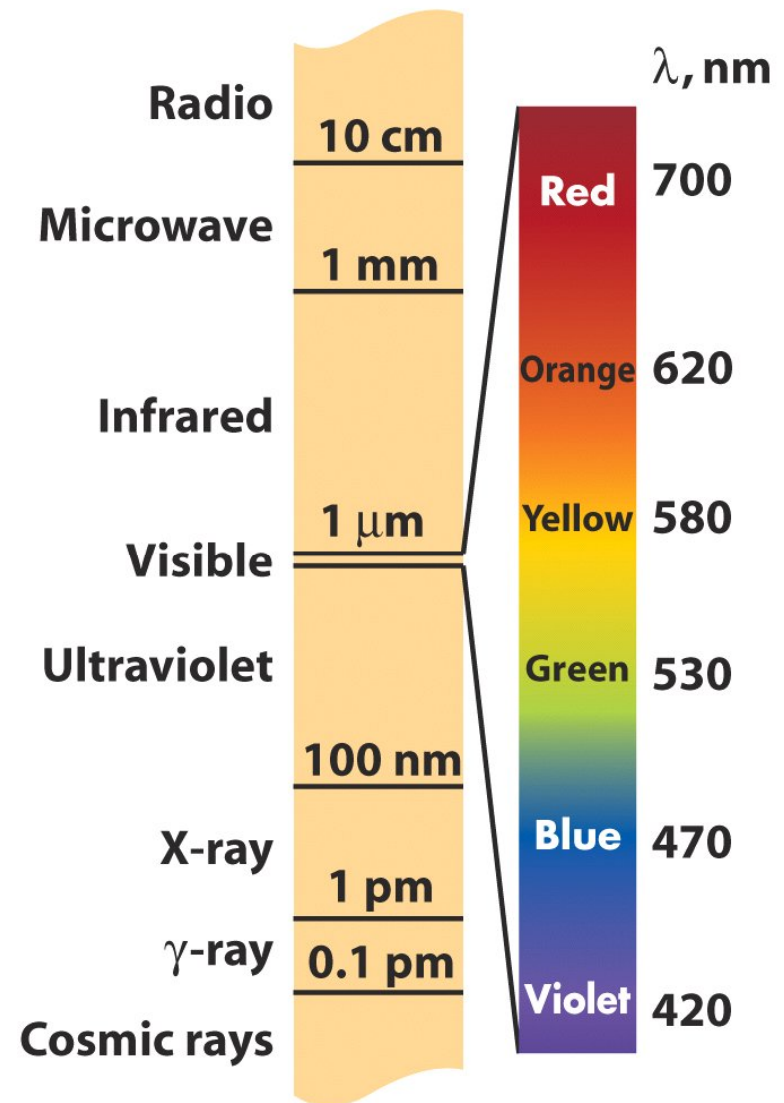
(Note significant figures please!)



$$c = \lambda \nu$$

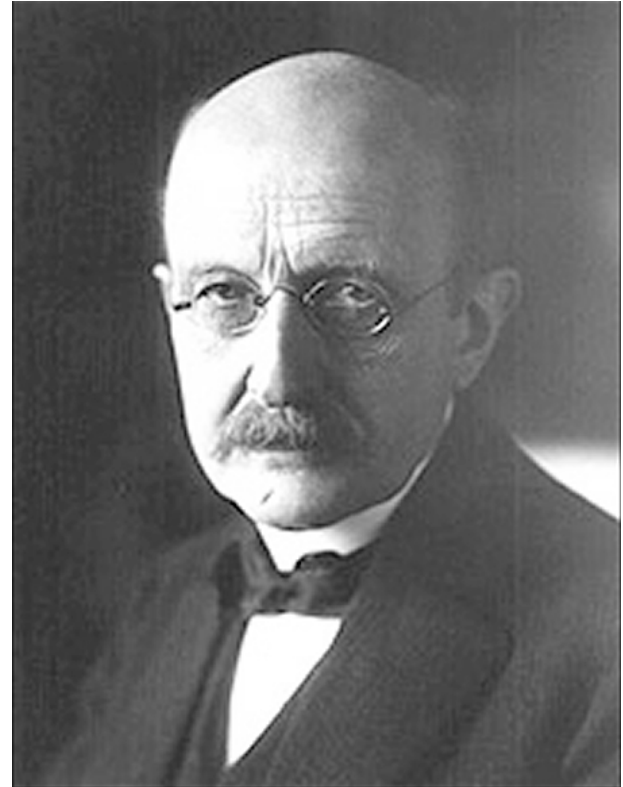
# The Electromagnetic Spectrum

- light travels with a velocity of  $3 \times 10^8$  m/s in vacuum.
- visible light ranges from 400 nm (violet) to 750 nm (red).
- Radio waves have lengths of cm - this is the reason antennae must be cm in length;
- X-rays have lengths of Angstroms - this is the reason we can use x-rays to elucidate the structure of matter composed of atoms (which are Angstroms in length).  
 **$1 \text{ \AA} = 0.1 \text{ nm} = 10^{-10} \text{ m}$**



# The Death of Classical Physics

"Experiments are the only means of knowledge at our disposal. The rest is poetry, imagination."



Max Planck  
1858-1947

# The Death of Classical Physics

Light energy is quantized.

$$E = h\nu = \frac{hc}{\lambda}$$

$h$  = Planck's constant

$6.626 \times 10^{-34}$  Joule sec

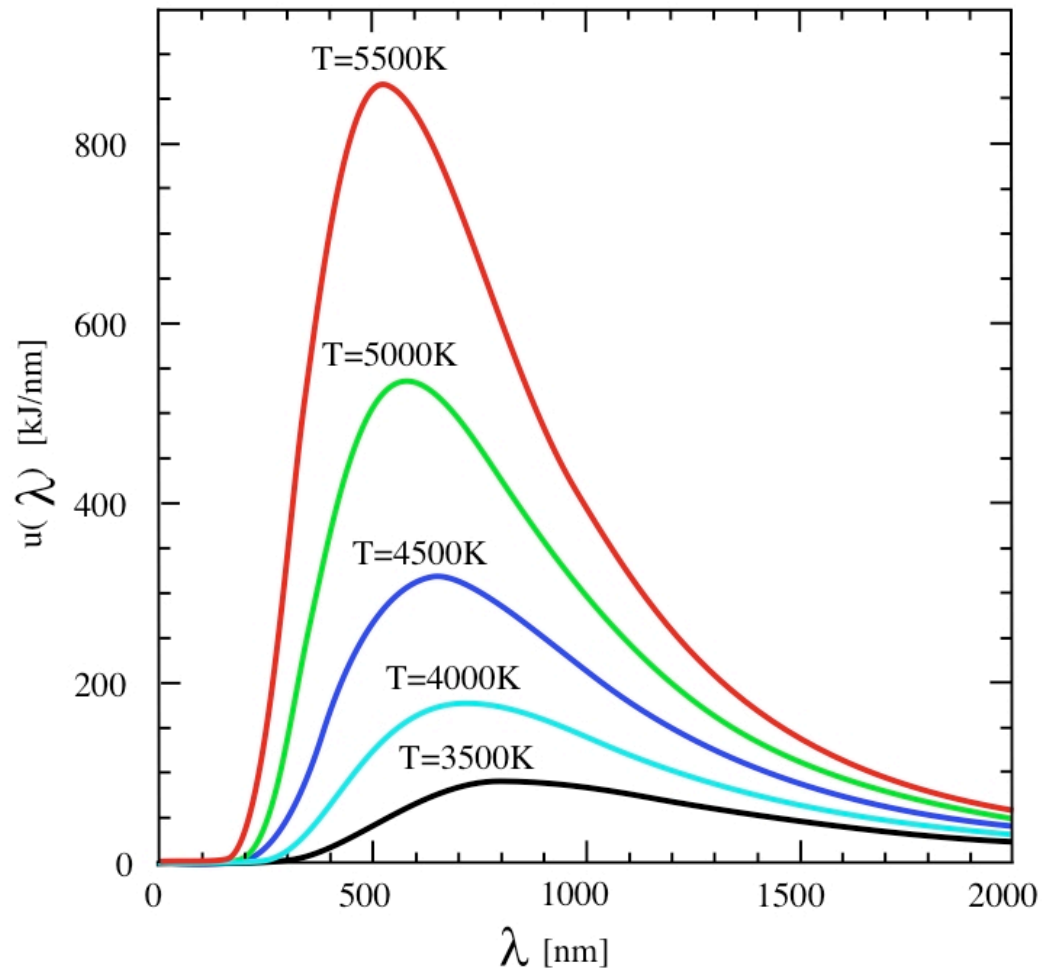


Max Planck  
1858-1947



# Blackbody Radiation

- Objects radiate light that has a spectrum based purely on the temperature of the object. We call this “blackbody radiation”.
- The intensity increases with  $T$ , and the peak intensity is observed at shorter and shorter wavelengths as  $T$  increases. The threshold for light emission in  $\lambda$  also increases with  $T$ .
- The classical theories were insufficient to predict the intensity versus wavelength seen in the plot at right.



[http://en.wikipedia.org/wiki/Image:Wiens\\_law.svg](http://en.wikipedia.org/wiki/Image:Wiens_law.svg)

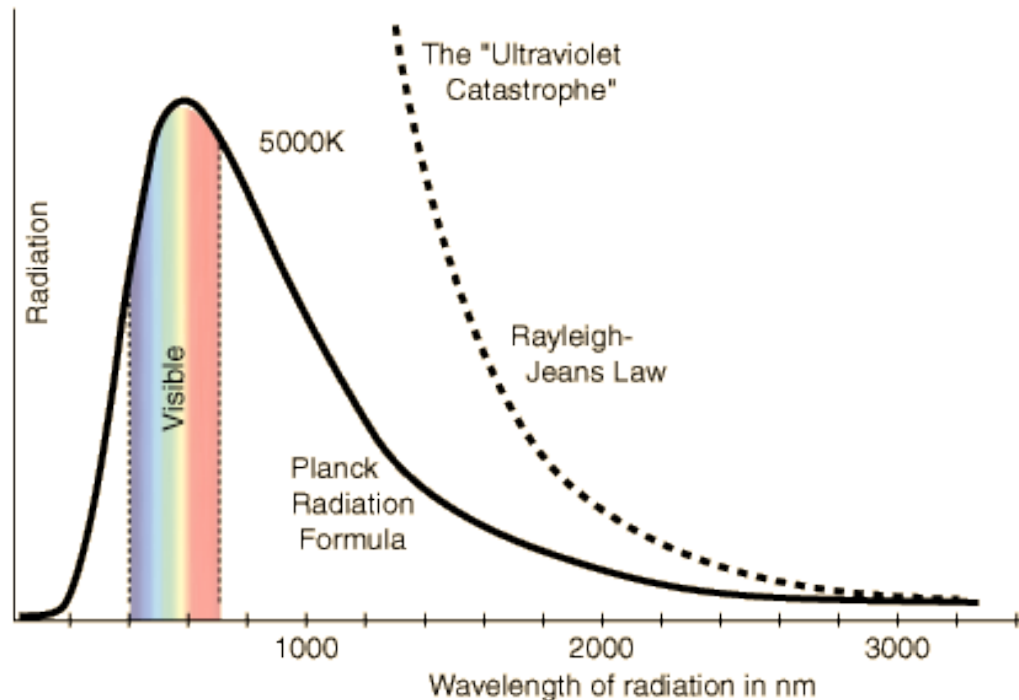
# Blackbody Radiation

## The “Ultraviolet Catastrophe”...

- The Rayleigh-Jeans law got the spectrum right at low  $\nu$ , but it blew up as  $\nu$  got larger.

**wrong**

$$\rho_{\nu}(T)d\nu = \frac{8\pi k_B T}{c^3} \nu^2 d\nu$$



<http://hyperphysics.phy-astr.gsu.edu/hbase/mod6.html>

# Blackbody Radiation

## The “Ultraviolet Catastrophe”...

- The Rayleigh-Jeans law got the spectrum right at low  $\nu$ , but it blew up as  $\nu$  got larger.
- Planck derived an equation based on a radical assumption: That the “generators” within the hot object could produce energy only in “packets” or “quanta” of  $h\nu$ , where  $h$  was a fudge factor.
- Planck adjusted the value of  $h$  in his equation to fit the data at one temperature, and he found that, in fact, it then fit the data at all temperatures.

wrong

$$\rho_{\nu}(T)d\nu = \frac{8\pi k_B T}{c^3} \nu^2 d\nu$$

right

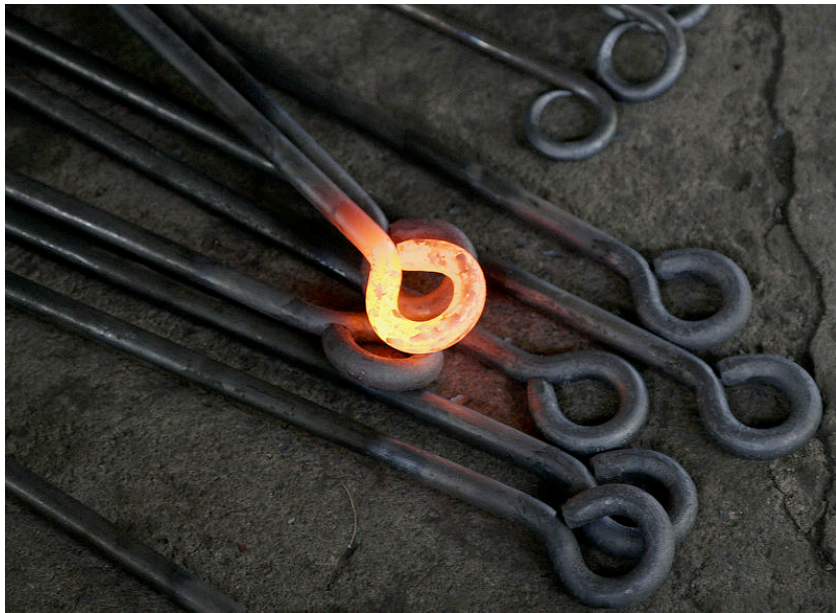
$$\rho_{\nu}(T)d\nu = \frac{8\pi h}{c^3} \frac{\nu^3 d\nu}{\exp\left(\frac{h\nu}{k_B T}\right) - 1}$$

# Blackbody Radiation

By differentiation, we can find the wavelength of maximum emission from a blackbody radiator. This is called:

$$\lambda_{\max} = \frac{2898.}{T}$$

Wien's Displacement Law

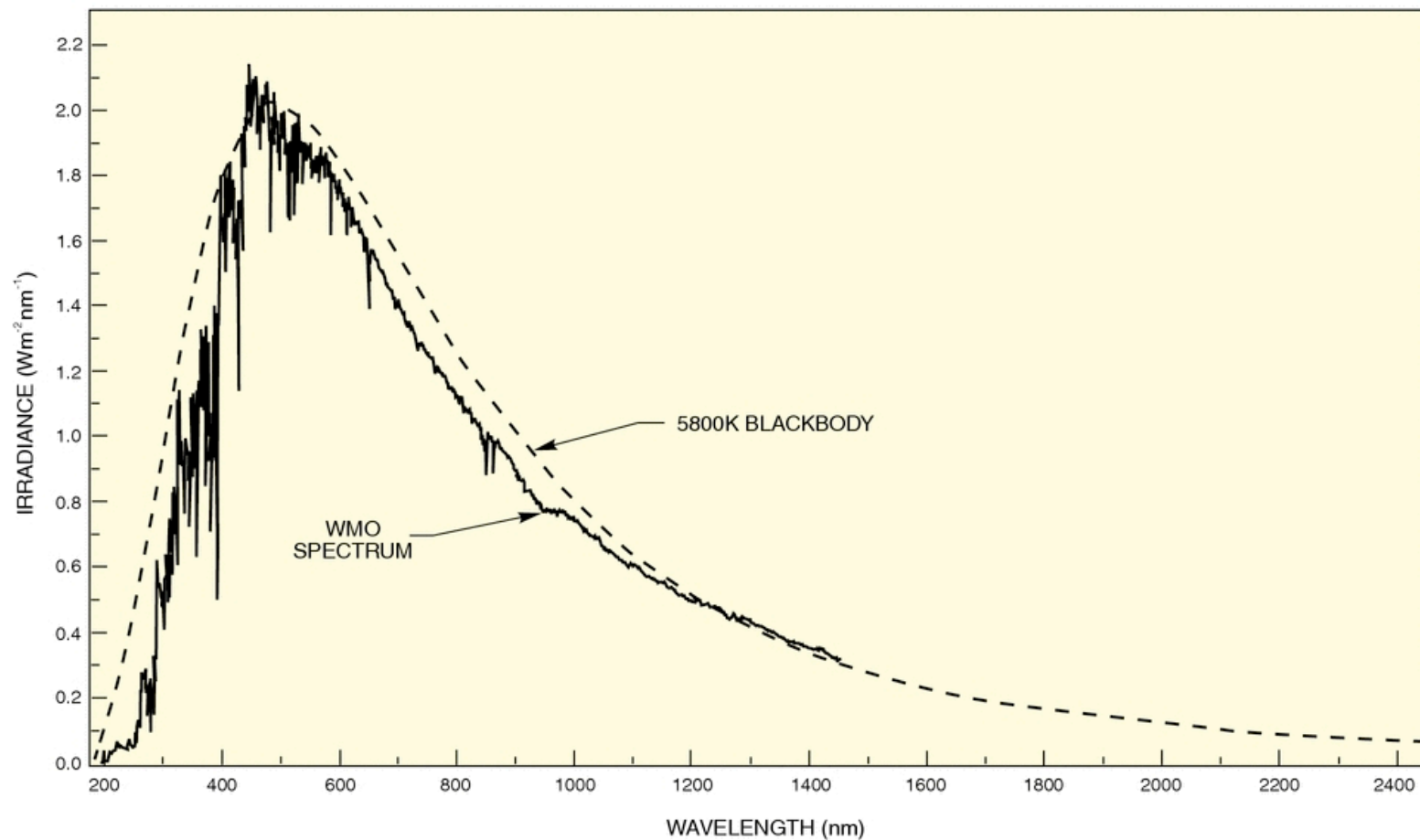


T is in Kelvin and Lambda is in microns in this equation. In SI units, the constant becomes:

$$2.898 \times 10^{-3} \text{ m K}$$

# Blackbody Radiation

The spectrum of the sun (outside the atmosphere) has the shape of a blackbody radiator with a maximum irradiance at a wavelength of 500 nm. This corresponds to a temperature of 5800K.



# Blackbody Radiation

Humans are blackbody radiators that emit light in the infrared.



$$\lambda_{\max} = \frac{2898.}{T}$$

$T = 37.0 \text{ C (310.K)}$  corresponds to a  $\lambda_{\max} = 9.35 \text{ microns}$ .

# Blackbody Radiation

Humans are blackbody radiators that emit light in the infrared.



Ear thermometers measure the blackbody emission of infrared light from the eardrum to get a temperature.

$T = 37.0\text{ C (310.K)}$  corresponds to a  
 $\lambda_{\text{max}} = 9.35\text{ microns.}$

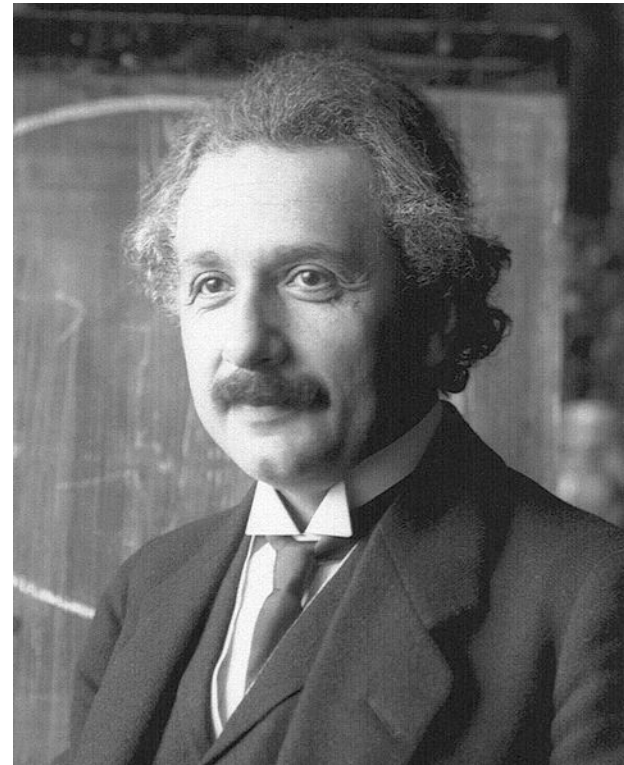




# Photons and Electrons



Heinrich Hertz  
1857-1894

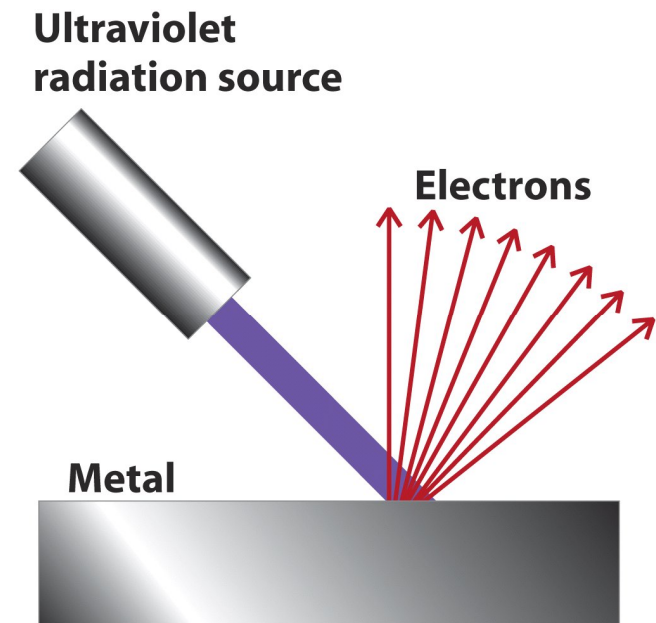


Albert Einstein  
1879-1955

# The Photoelectric Effect (1887)

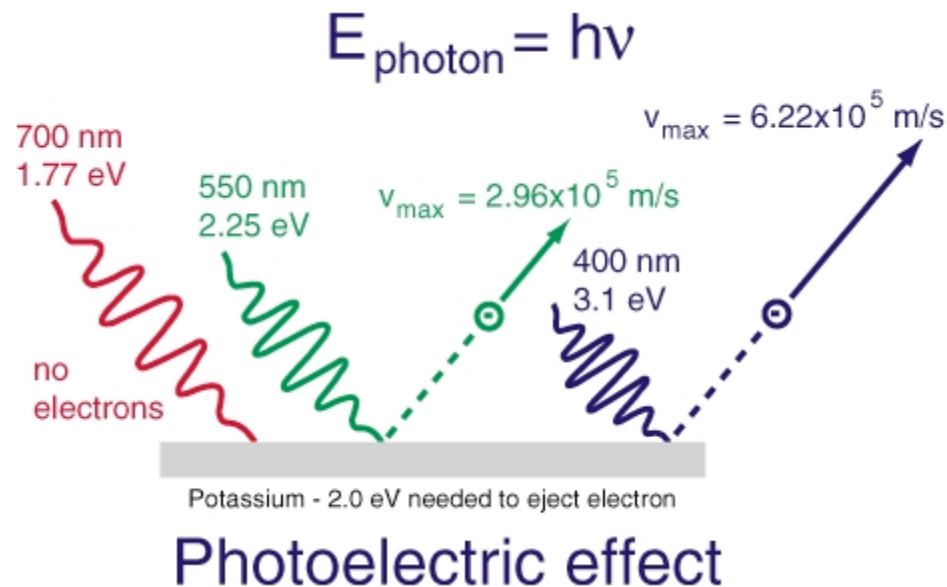


Heinrich Hertz  
1857-1894



aka "The Hertz Effect"

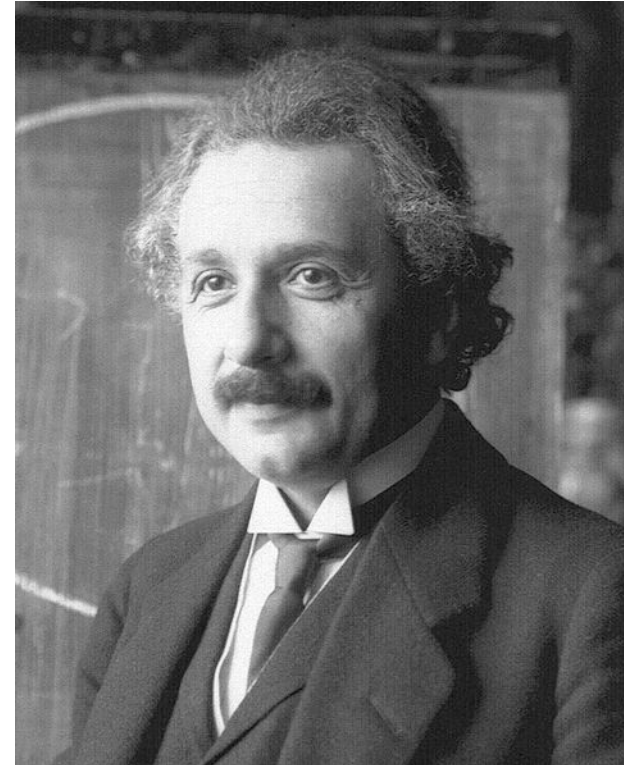
# The Photoelectric Effect (1905)



$$h\nu = E_{\text{kinetic}} + \phi_{\text{metal}}$$

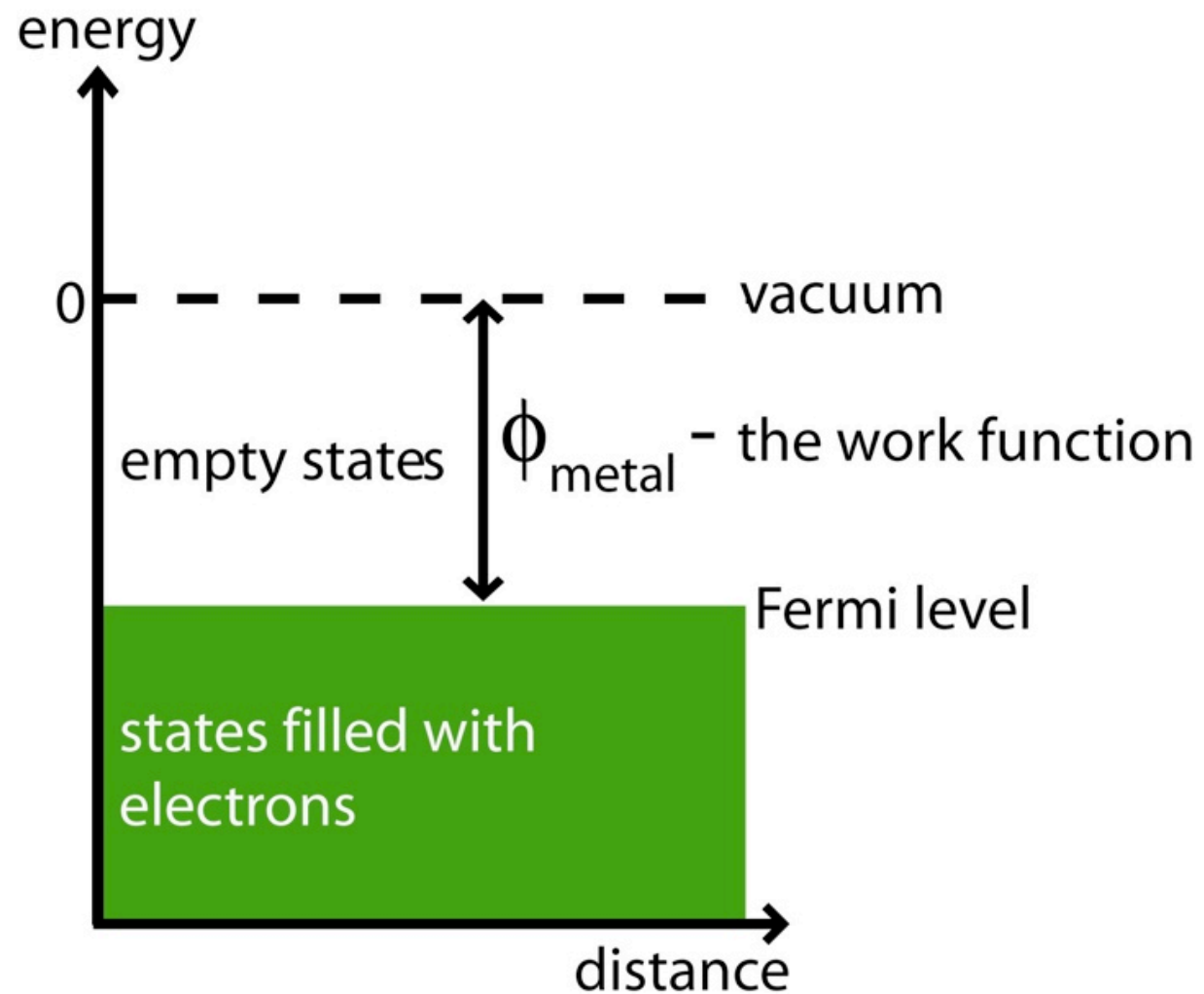
Wavelength dependence  
must be explained by the  
existence of quantized light.

Nobel Prize (his only one) awarded in 1921.



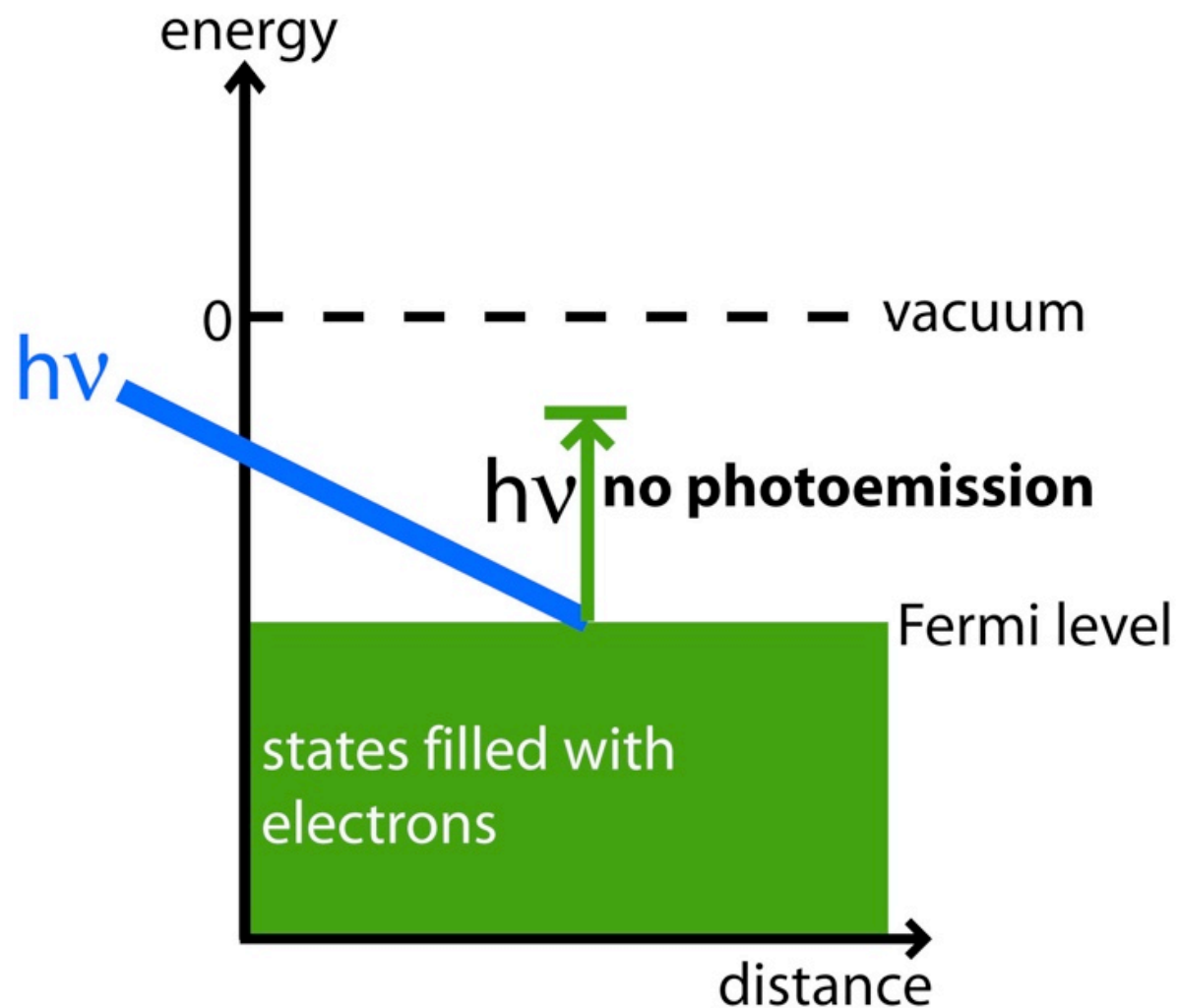
Albert Einstein  
1879-1955

how does photoemission work at metals?



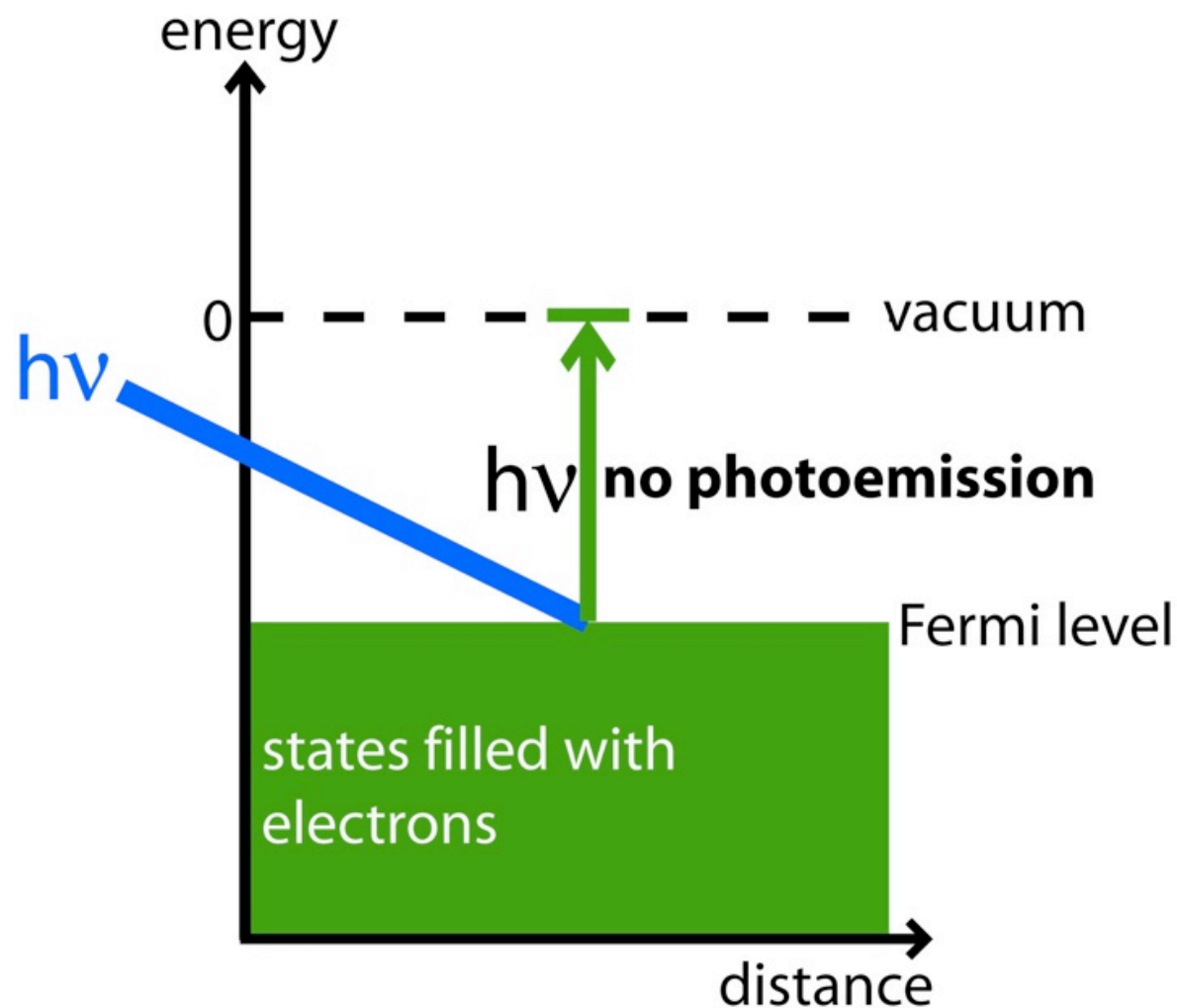
You shine light on a metal; **3 things can happen:**

Case 1:  $h\nu < \phi_{\text{metal}}$



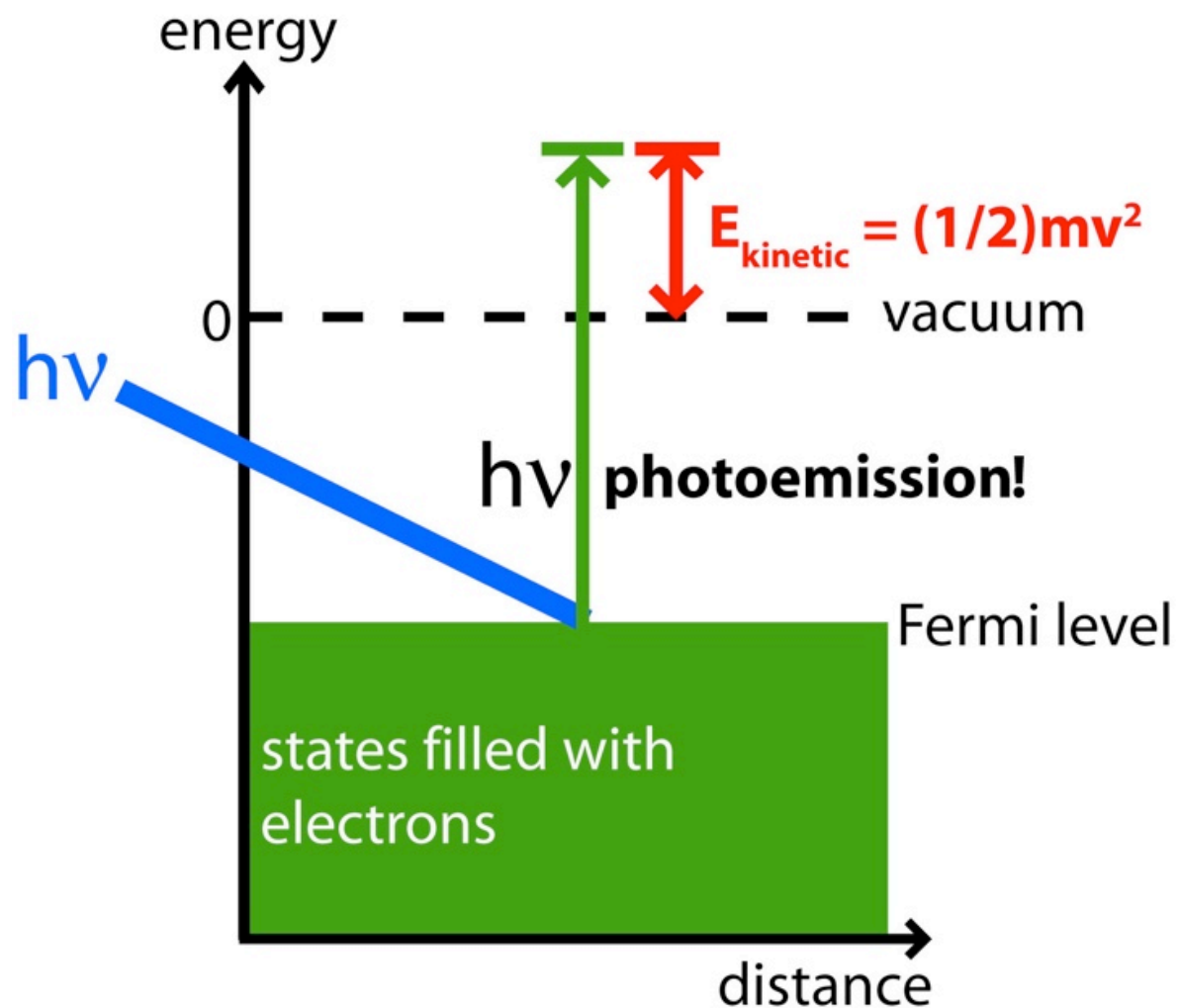
You shine light on a metal; **3 things can happen:**

Case 2:  $h\nu = \phi_{\text{metal}}$



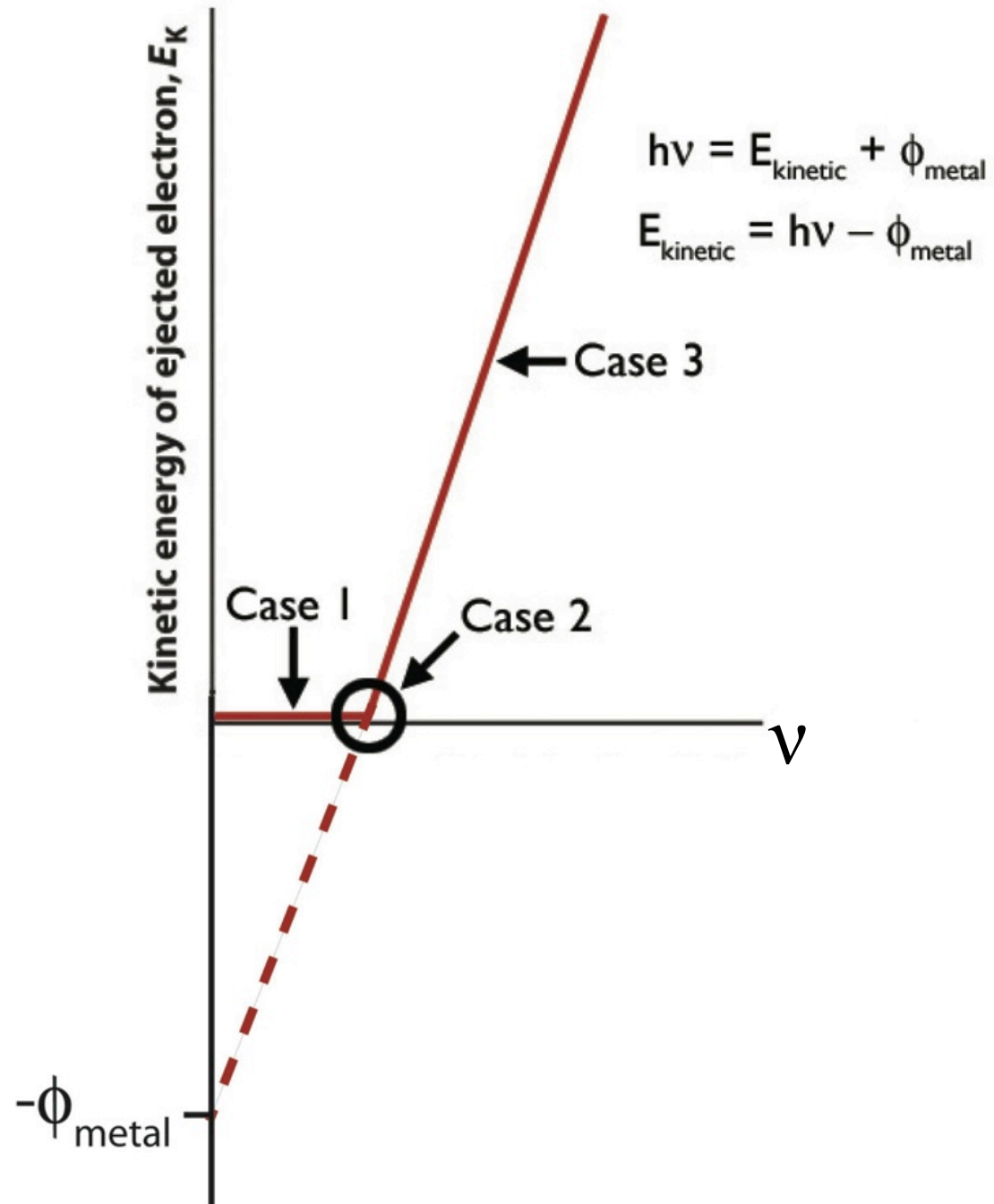
You shine light on a metal; **3 things can happen:**

Case 3:  $h\nu > \phi_{\text{metal}}$





what do we measure in the laboratory?



*Actually...*

## The origin of the word "photon"

"I therefore take the liberty of proposing for this hypothetical new atom, which is not light but plays an essential part in every process of radiation, the name *photon*."

-Gilbert N. Lewis, 1926

Although Planck and Einstein advanced the concept of quanta, Einstein did not use the word photon in his early writings and as far as my reading goes, he never did. The word "photon" originated from Gilbert N. Lewis years after Einstein's photoelectric paper and appeared in a letter to the editor of Nature magazine (Vol. 118, Part 2, December 18, 1926, page 874-875).

<http://www.nobeliefs.com/photon.htm>

# The Death of Classical Physics

Light energy is quantized.

$$E = h\nu = \frac{hc}{\lambda}$$

$h$  = Planck's constant

$6.626 \times 10^{-34}$  Joule sec



Max Planck  
1858-1947

**Question:** Calculate the energy (in eV) of a “typical” red photon with a wavelength of 600.0 nm. (eV = electron volt)

Key equations:  $c = \lambda\nu$  and  $E = h\nu$

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Step 1 - Convert  $\lambda$  into  $\nu$ :

$$c = 2.998 \times 10^8 \text{ m/s} = \lambda\nu$$

$$\nu = (2.998 \times 10^8 \text{ m/s}) / (600.0 \times 10^{-9} \text{ m})$$

$$4.997 \times 10^{14} \text{ s}^{-1} \text{ or } 4.997 \times 10^{14} \text{ Hz.}$$

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Step 2 - Multiply by  $h$ :

$$E = (6.626 \times 10^{-34} \text{ J s})(4.997 \times 10^{14} \text{ s}^{-1})$$

$$E = 3.311 \times 10^{-19} \text{ J}$$

**Question:** Calculate the energy (in eV) of a “typical” red photon with a wavelength of 600.0 nm. (eV = electron volt)

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$$E = (6.626 \times 10^{-34} \text{ J s})(4.997 \times 10^{14} \text{ s}^{-1})$$

$$E = 3.311 \times 10^{-19} \text{ J}$$

Step 3 - Convert to eV:  $1 \text{ eV} = 1.602 \times 10^{-19} \text{ J/eV}$

$$E = (3.311 \times 10^{-19} \text{ J}) / (1.602 \times 10^{-19} \text{ J/eV})$$

$$E = 2.067 \text{ eV}$$



**Question:** Calculate the energy (in eV) of a “typical” red photon with a wavelength of 600.0 nm. (eV = electron volt)

Key equations:  $E = hc/\lambda$

Shortcut:

$$c = 2.998 \times 10^8 \text{ m/s}$$

$$h = (6.626 \times 10^{-34} \text{ J s}) / (1.602 \times 10^{-19} \text{ J/eV})$$

$$E = (1.240 \times 10^{-6}) / (\lambda \text{ in m})$$

$$E = 1240. / (\lambda \text{ in nm})$$

Calculation:

$$E = 1240. / 600.0$$

$$E = 2.067 \text{ eV}$$

# The Death of Classical Physics

from Planck, we have:

- the birth of quantum theory.
- $h$  - Planck's constant.  $6.626 \times 10^{-34} \text{ J s}$ .
- the quantization of energy -  
a concept underlying all of  
quantum mechanics.

We will use quantum mechanics to  
understand Chemistry!



Max Planck  
1858-1947

# De Broglie Wavelength (1924)

For photons:

Special Relativity  
and Planck

momentum  
of a photon

$$E = pc = \frac{hc}{\lambda}$$

$$p = \frac{h}{\lambda}$$



Louis de Broglie  
1892-1987

# De Broglie Wavelength (1924)

For photons:

Special Relativity  
and Planck

momentum  
of a photon

wavelength  
of a particle

$$E = pc = \frac{hc}{\lambda}$$

$$p = \frac{h}{\lambda}$$

$$\lambda = \frac{h}{p}$$



why not ALL particles?



Louis de Broglie  
1892-1987

# De Broglie Wavelength (1924)

de Broglie wavelength  $\longrightarrow$

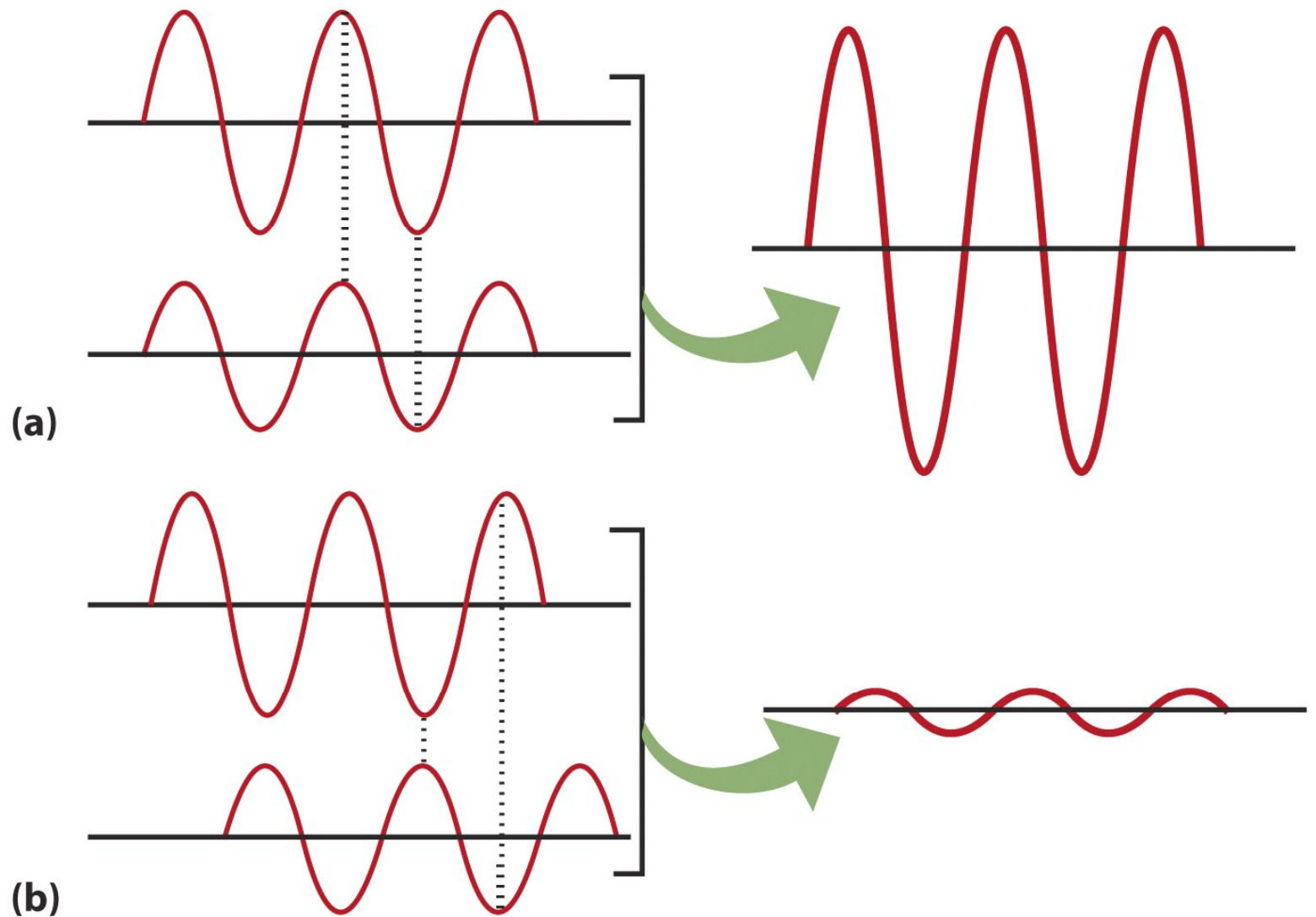
$$\lambda = \frac{h}{p}$$

$6.626 \times 10^{-34} \text{ J s}$   
 $\downarrow$   
 $h$   
 $\uparrow$   
 $p$   
mass x velocity



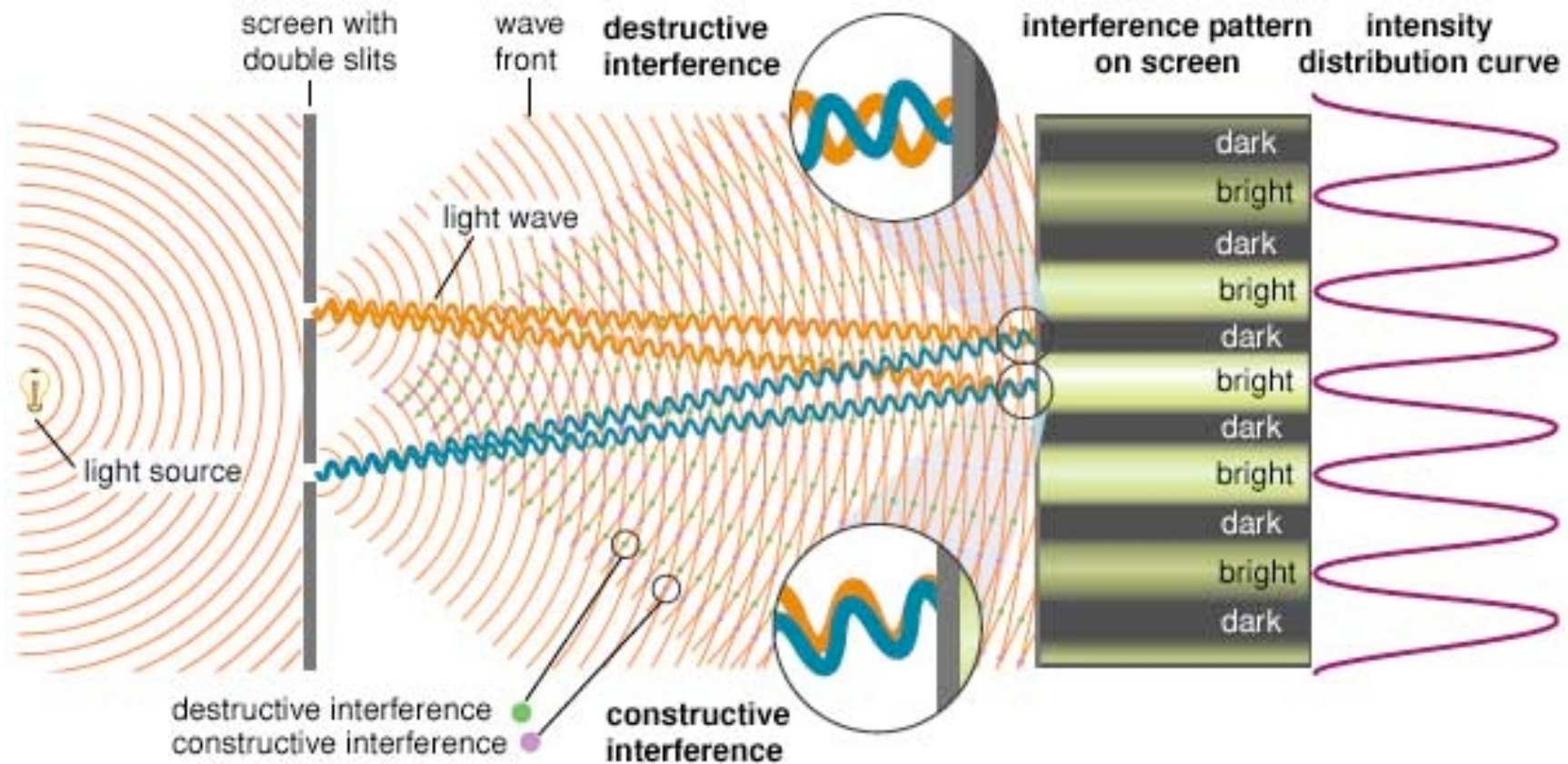
Louis de Broglie  
1892-1987

A signature property of waves is the phenomenon of interference: Waves can interfere either constructively (a) or destructively (b)...





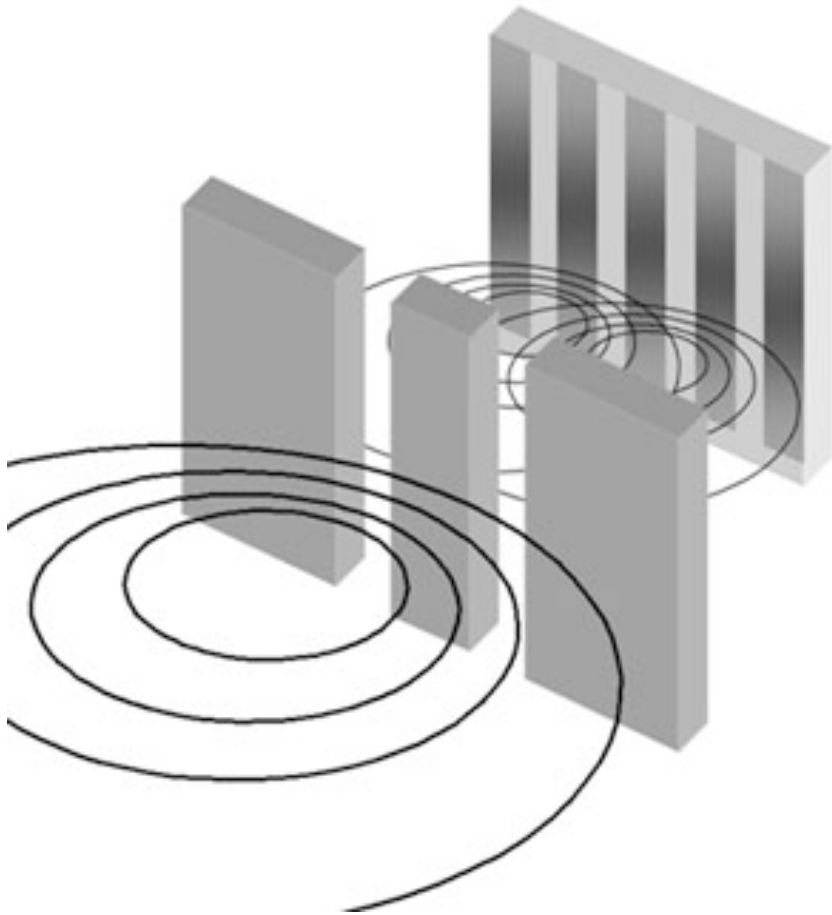
## Thomas Young's two-slit experiment with light (1802).



Young's experiment convinced physicists that light was a wave.

# Light diffracts like waves:

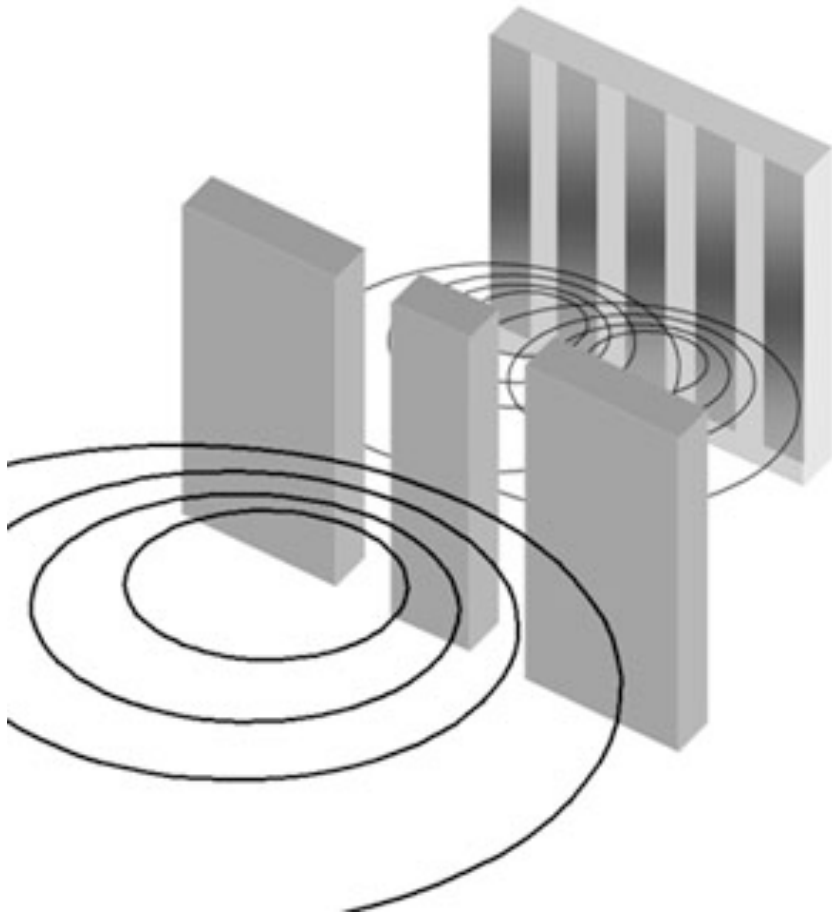
Thomas Young's two-slit experiment with light (1802).



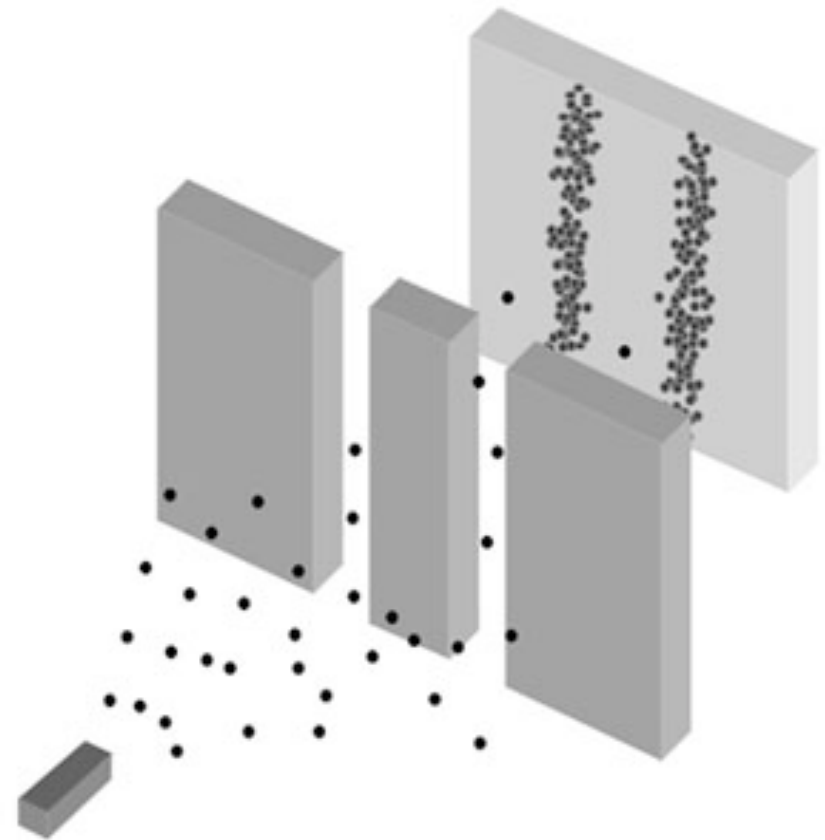


# Electrons also diffract like waves!

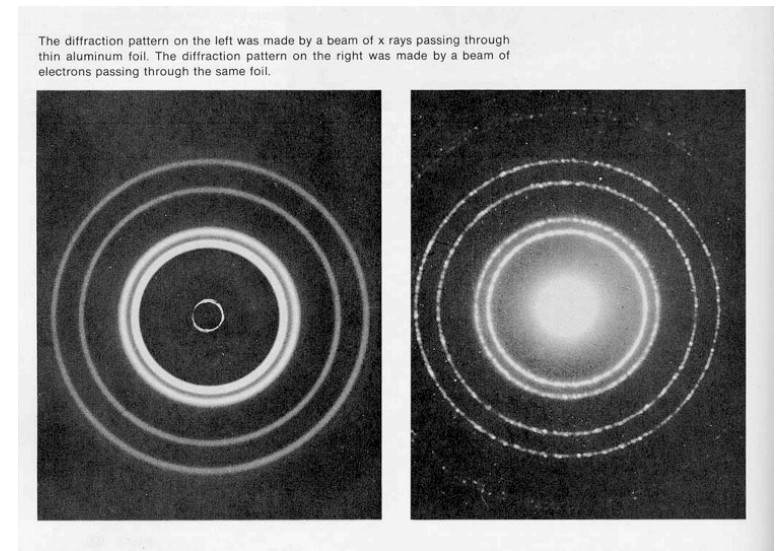
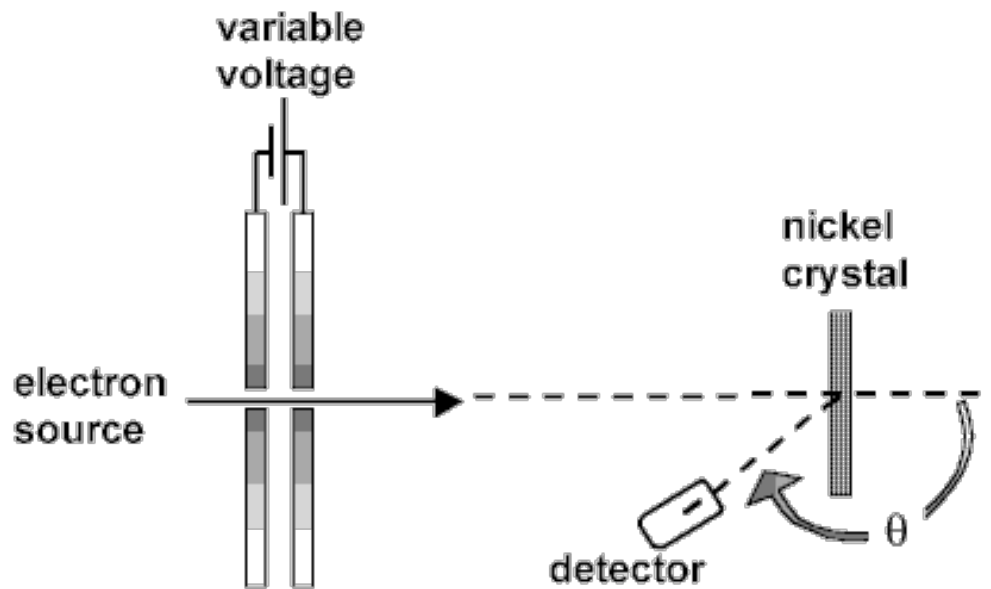
Thomas Young's two-slit experiment with light (1802).



Davisson and Germer electron diffraction (1928).



Davission and Germer measured electron diffraction from a Nickel crystal in 1928.



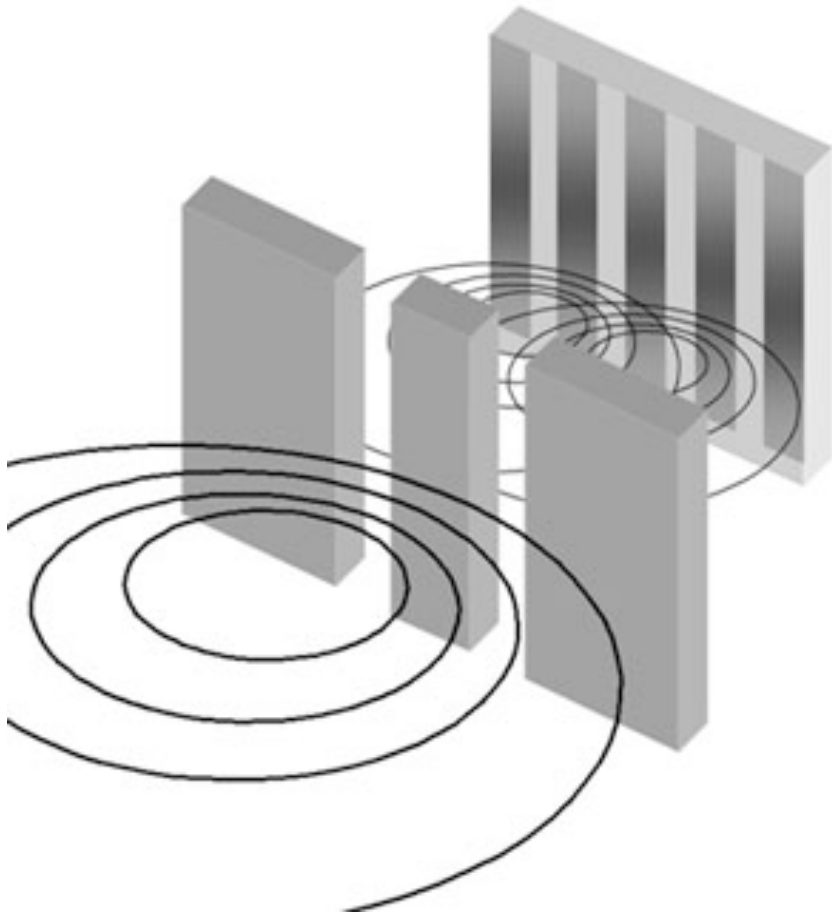
X-rays

electrons

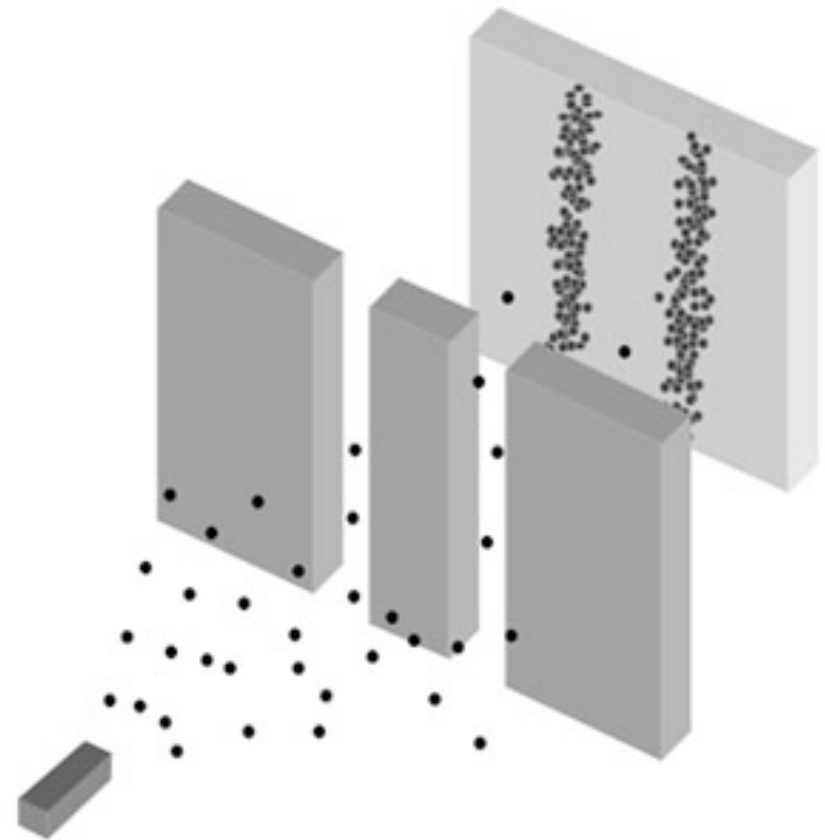
The de Broglie wavelength of electrons in the lab is *Angstroms* ( $10^{-10} \text{ m}$ ), so the “slits” have to be atomic scale - Davission and Germer used a Nickel crystal.

# Electrons also diffract like waves!

Thomas Young's two-slit experiment with light (1802).



Davisson and Germer electron diffraction (1928).

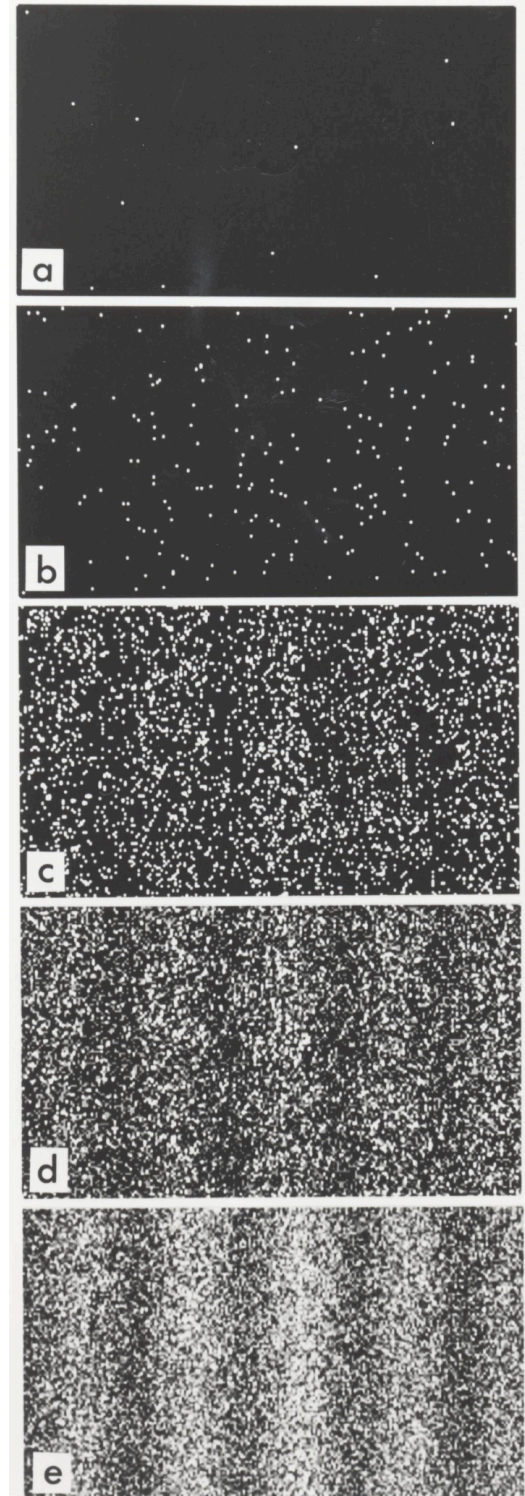


Question: How many electrons do we need to see interference?

Answer: one!

Modern version of Thomas Young's two-slit experiment performed with electrons.

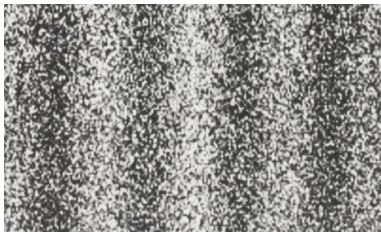
*American Journal of Physics* in 1974 (volume 42, pp4-11)



# De Broglie Wavelength (1924)

*All particles have a wave nature, just as electromagnetic waves have a *particle* nature.*

$$\lambda = \frac{h}{p}$$



Electron diffraction is observed, even when the flux of electrons is so low, that *consecutive electrons encounter these slits one at a time.*



Louis de Broglie  
1892-1987

Clearly, we need a way to predict the bizarre behavior of these particles.  
Enter the New Wave Stars: Schrödinger, Heisenburg and Dirac!



# The Birth of Quantum Mechanics



Louis de Broglie  
1892-1987



Erwin Schrödinger  
1887-1961



Paul Dirac  
1902-1984



Werner Heisenberg  
1901-1976

## New Wave Rock Stars